P2SC-ROB-WR-761 - 20241021	P2SC Weekly report	* **** <u>****</u>
Period covered: Date:	,	Royal Observatory of Belgium -
	Laurence Wauters Marie Dominique	PROBA2 Science Center
To:	LYRA PI, marie.dominique@sidc.be SWAP PI, elke.dhuys@sidc.be	https://proba2.sidc.be ++ 32 (0) 2 3730559
CC:	ROB DIR, ronald@oma.be ESA Redu, Rene.Wittmann@esa.int and Marcus.De.Deus.Silva@esa.int ESA D/SRE, Joe.Zender@esa.int ESA D/TEC, Juha-Pekka.Luntama@esa.int and Melanie.Heil@esa.int	

# 1. Science

# Solar & Space weather events

The level of solar activity<sup>1</sup> fluctuated between **low and high** this week.

Only M- and X-flares are mentioned, the most energetic one(s) per day are presented in **bold**:

	Monday 21 Oct	Tuesday 22 Oct	Wednesday 23 Oct	Thursday 24 Oct	Friday 25 Oct	Saturday 26 Oct	Sunday 27 Oct
Activity	low	low	low	high	moderate	high	moderate
Flares	-	-	-	M1.2, X3.3	M1.1	M1.6, M2.8,M2.2, X1.8, M1.5	M2.8

<sup>&</sup>lt;sup>1</sup> See appendix. All timings are given in UT.

#### Solar Activity

Solar flare activity fluctuated from low to high during the week.

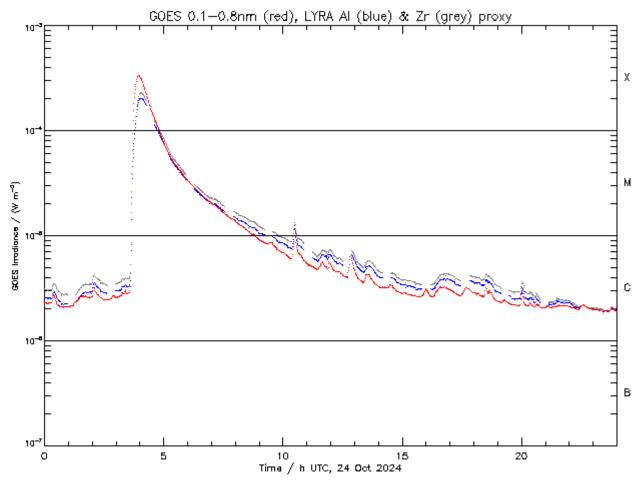
In order to view the activity of this week in more detail, we suggest to go to the following website from which all the daily (normal and difference) movies can be accessed: <u>https://proba2.oma.be/ssa</u> This page also lists the recorded flaring events.

A weekly overview movie can be found here (SWAP week 761).

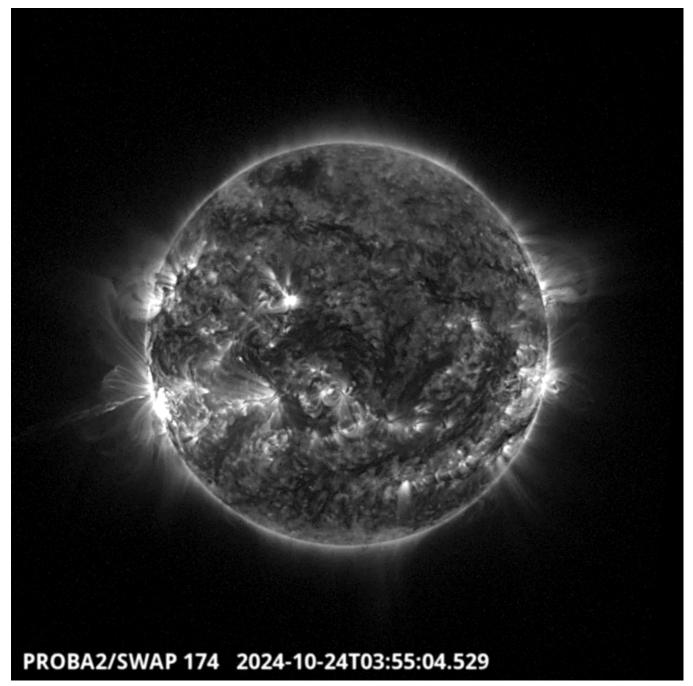
Details about some of this week's events can be found further below.

If any of the linked movies are unavailable they can be found in the P2SC movie repository here

# Thursday Oct 24

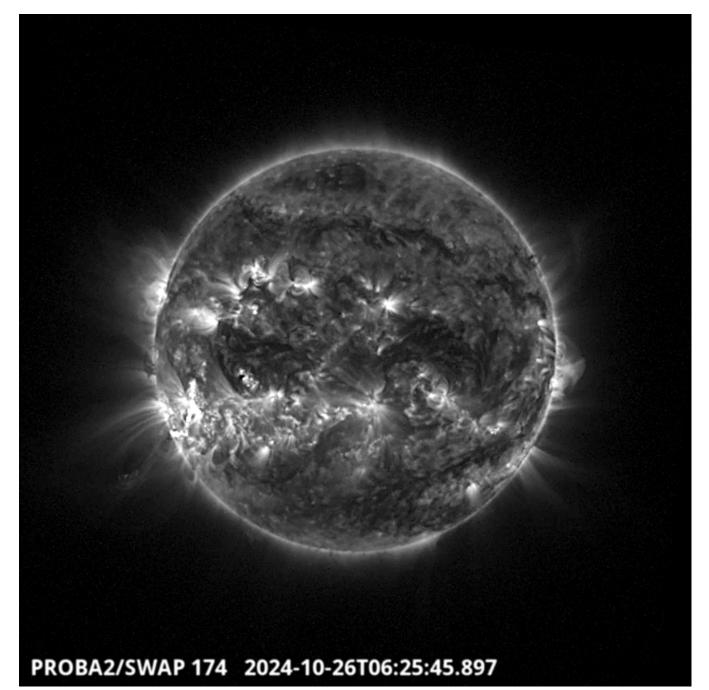


ROB/SIDC, Brussels, Belgium



The largest flare of the week is a X3.3 visible on the LYRA and SWAP image above. It was produced by the 3860 NOAA active region located on the South East limb of the solar disk. Find a SWAP movie of the event <u>here</u>.

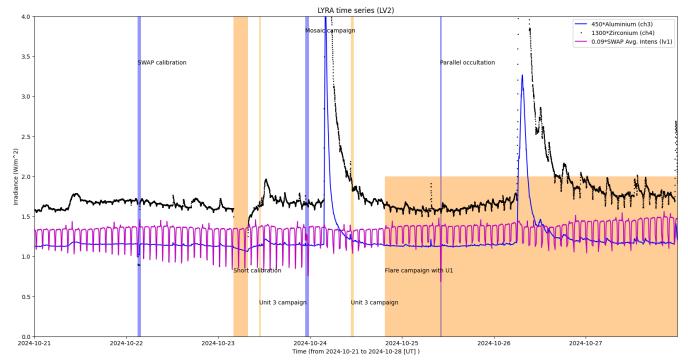
### Saturday Oct 26



A M9.5 flare erupted around 6:23 UT, it originated from the active region NOAA 3873. It was followed by another flare, a X1.8 peaking at 7:19 and issued from the same active region located on the South East part of the solar disk. Find a SWAP movie of the event <u>here</u>. An overview of the weekly LYRA & SWAP data is provided below:

The following curves are visible:

- black: Zirconium Channel LYRA Unit 2
- blue: Aluminium Channel of LYRA Unit 2
- purple: SWAVINT (SWAP Average Intensity; integrated solar intensity per SWAP image pixel )



## **Operations and Calibrations:**

The blue shaded periods related to SWAP, correspond to, from left to right:

- SWAP calibration, 2024-Oct-22
- Mosaïc campaign, 2024-Oct-23
- SWAP parallel occultation with LYRA, 2024-Oct-25.

The orange shaded periods related to LYRA correspond to, from left to right:

- LYRA short calibration, 2024-Oct-23
- Daily unit 3 campaigns during occultations, 2024-Oct-24
- LYRA flare campaign, since 2024-Oct-24 until 2024-Oct-27

The red shaded periods related to other issues corresponds to:

• None

# 2. LYRA instrument status

IOS

Start IOS	Mon Oct 21 2024	LYIOS01124
End IOS	Sun Oct 27 2024	LYIOS01125

## LYRA detector temperature

LYRA detector 2 temperature globally varied between 51.04 and 56.38 °C.

# 3. SWAP instrument status

## MCPM errors

The number of MCPM recoverable errors increased from 64878 to 65392. The number of MCPM unrecoverable errors remained at 3135.

#### IOS

Start IOS	Mon Oct 21 2024	IOS01228
End IOS	Sun Oct 27 2024	IOS01230

### SWAP detector temperature

The SWAP Cold Finger Temperature globally varied between 1.59 and 3.83 °C.

# 4. PROBA2 Science Center Status

The following changes were made to the P2SC:

• None.

# 5. Data reception & discussions with MOC

### Passes

The delivery of the passes for this week (passes 48579 to 48664) was nominal, except for:

• None.

#### Data coverage HK

All HK data files (LYRA\_AD) have been received, except:

• None.

### Data coverage SWAP

All SWAP Science data files (BINSWAP) have been received, except:

• None.

Total number of images between 2024 Oct 21 0UT and 2024 Oct 28 0UT: 4450 Highest cadence in this period: 17 seconds Average cadence in this period: 135.91 seconds Number of image gaps larger than 300 seconds: 162 Largest data gap: 17.37 minutes

## Data coverage LYRA

All LYRA Science data files (BINLYRA) have been received, except:

None

# 6. APPENDIX: Frequently used acronyms

ADPMS	Advanced Data and Power Management System
AOCS	Attitude and Orbit Control System
APS	Active Pixel image Sensor
ASIC	Application Specific Integrated Circuit
BBE	Base Band Equipment
CME	Coronal Mass Ejection
COGEX	Cool Gas Generator Experiment
CRC	Cyclic Redundancy Check
DAC	Data Acquisition Controller
DBR	Deployment, backup & recovery
DDA	Decommutated data archive
ESP	Experimental Solar Panel
FITS	Flexible Image Transport System
FOV	Field Of View FPA Focal Plane Assembly
FPGA	Field Programmable Gate Arrays
GPS	Global Positioning System
НК	Housekeeping
IOS	Instrument Operations Sheet
LED	Light Emitting Diode
LYRA	LYman alpha RAdiometer
LYTMR	LYRA Telemetry Reformatter (software module of P2SC)
LYEDG	LYRA Engineering Data Generator (software module of P2SC)
MCPM	Mass Memory, Compression and Packetisation Module
MOC	Mission Operation Center
NDR	Non Destructive Readout
OBSW	On board Software
PI	Principal Investigator
P2SC	PROBA2 Science Center
ROB	Royal Observatory of Belgium
SAA	South Atlantic Anomaly
SEU	Single Event Upset
SoFAST	Solar Feature Automated Search Tool
SWAP	Sun Watcher using APS detector and image Processing
SWAVINT	SWAP AVerage INTensity
SWBSDG	SWAP Base Science Data Generator
SWEDG	SWAP Engineering Data Generator (software module of P2SC)
SWTMR	SWAP Telemetry Reformatter (software module of P2SC)
TBC	To Be Confirmed
TBD	To Be Defined
TC	Telecommand
UTC	Coordinated Universal Time
UV	Ultraviolet
VFC	Voltage to Frequency Converter

# 7. APPENDIX Solar Activity Definitions

In the science section we use the following solar activity standards.

The standard scale for solar activity is:

- very low (almost no flares, only B)
- low (a few C flares)
- moderate (many C flares and at least an M flare)
- high (several M flares and an X flare)
- very high (continuous background of C flares, numerous M flares, more than one X flare)