P2SC-ROB-WR-678 - 20230320	P2SC Weekly report	**** ****
Period covered: Date:	′	Royal Observatory of Belgium
Written by: Approved by:		PROBA2 Science Center
То:	LYRA PI, marie.dominique@sidc.be SWAP PI, elke.dhuys@sidc.be	https://proba2.sidc.be ++ 32 (0) 2 3730559
CC:	ROB DIR, ronald@oma.be ESA Redu, Rene.Wittmann@esa.int and Marcus.De.Deus.Silva@esa.int ESA D/SRE, Joe.Zender@esa.int ESA D/TEC, Juha-Pekka.Luntama@esa.int and Melanie.Heil@esa.int	

## 1. Science

## Solar & Space weather events

The level of solar activity<sup>1</sup> fluctuated between **very low and moderate** this week.

Only M- and X-flares are mentioned, the most energetic one(s) per day are presented in **bold**:

	Monday 20 Mar	Tuesday 21 Mar	Wednesday 22 Mar	Thursday 23 Mar	Friday 24 Mar	Saturday 25 Mar	Sunday 26 Mar
Activity	moderate	low	low	very low	low	low	low
Flares	M1.2	-	-	-	-	-	-

<sup>&</sup>lt;sup>1</sup> See appendix. All timings are given in UT.

#### **Solar Activity**

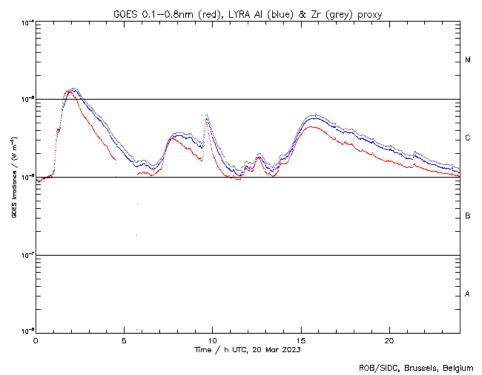
Solar flare activity fluctuated from very low to moderate during the week. In order to view the activity of this week in more detail, we suggest to go to the following website from which all the daily (normal and difference) movies can be accessed: <a href="https://proba2.oma.be/ssa">https://proba2.oma.be/ssa</a>
This page also lists the recorded flaring events.

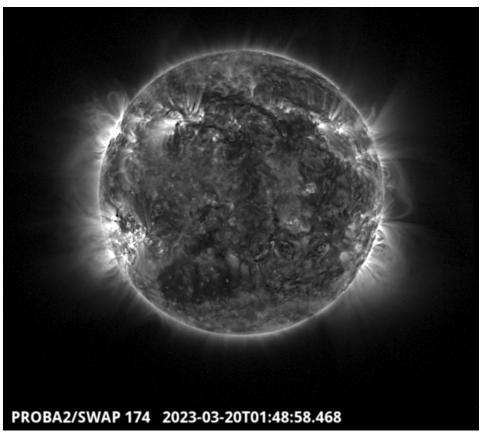
A weekly overview movie can be found here (SWAP week 678).

Details about some of this week's events can be found further below.

If any of the linked movies are unavailable they can be found in the P2SC movie repository <a href="here">here</a>

#### Monday Mar 20





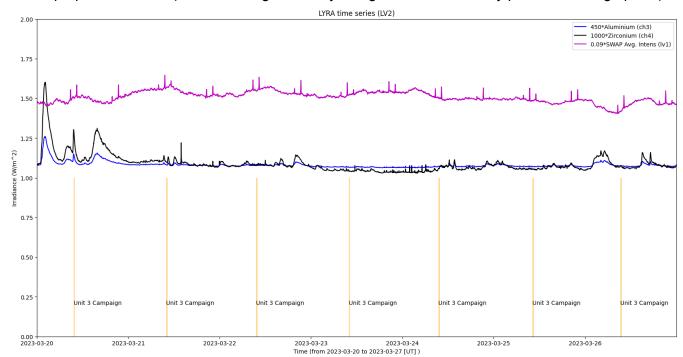
The largest and only M-flare of the week, an M1.2, was observed by LYRA (top panel) and SWAP (bottom panel). The flare occurred on 2023-Mar-20 (peak at 01:48 UT) in the south-eastern hemisphere, and it was associated with NOAA AR3256.

Find a SWAP movie of the event here.

An overview of the weekly LYRA & SWAP data is provided below:

The following curves are visible:

- black: Zirconium Channel LYRA Unit 2
- blue: Aluminium Channel of LYRA Unit 2
- purple: SWAVINT (SWAP Average Intensity; integrated solar intensity per SWAP image pixel)



#### **Operations and Calibrations:**

The blue shaded periods related to SWAP, correspond to, from left to right:

None

The orange shaded periods related to LYRA correspond to, from left to right:

- Daily Unit 3 campaign, 2023-Mar-20
- Daily Unit 3 campaign, 2023-Mar-21
- Daily Unit 3 campaign, 2023-Mar-22
- Daily Unit 3 campaign, 2023-Mar-23
- Daily Unit 3 campaign, 2023-Mar-24
- Daily Unit 3 campaign, 2023-Mar-25
- Daily Unit 3 campaign, 2023-Mar-26

The red shaded periods related to other issues corresponds to:

None

# 2. LYRA instrument status

#### IOS

Start IOS	Mon Mar 20 2023	LYIOS00997
End IOS	Sun Mar 26 2023	LYIOS00998

## LYRA detector temperature

LYRA detector 2 temperature globally varied between 51.3 and 52.15  $^{\circ}\text{C}.$ 

## 3. SWAP instrument status

#### **MCPM** errors

The number of MCPM recoverable errors increased from 39583 to 39876.

The number of MCPM unrecoverable errors remained at 3135.

#### IOS

Start IOS	Mon Mar 20 2023	IOS001111
End IOS	Sun Mar 26 2023	IOS001112

## **SWAP** detector temperature

The SWAP Cold Finger Temperature globally varied between 0.15 and 1.11 °C.

# 4. PROBA2 Science Center Status

The following changes were made to the P2SC:

• None.

## 5. Data reception & discussions with MOC

#### **Passes**

The delivery of the passes for this week (passes 43560 to 43619) was nominal, except for:

 due to file transfer issues caused by network slowdown in Redu, the files for passes 43574 -43624 have been fully received on March 27 and gaps have been filled.

#### Data coverage HK

All HK data files (LYRA\_AD) have been received, except:

None.

#### **Data coverage SWAP**

All SWAP Science data files (BINSWAP) have been received, except:

None.

Total number of images between 2023 Mar 20 0UT and 2023 Mar 27 0UT: 4354

Highest cadence in this period: 0 seconds

Average cadence in this period: 138.91 seconds Number of image gaps larger than 300 seconds: 209

Largest data gap: 11.00 minutes

#### **Data coverage LYRA**

All LYRA Science data files (BINLYRA) have been received, except:

None

### 6. APPENDIX: Frequently used acronyms

ADPMS Advanced Data and Power Management System

AOCS Attitude and Orbit Control System

APS Active Pixel image Sensor

ASIC Application Specific Integrated Circuit

BBE Base Band Equipment
CME Coronal Mass Ejection

COGEX Cool Gas Generator Experiment

CRC Cyclic Redundancy Check
DAC Data Acquisition Controller
DBR Deployment, backup & recovery
DDA Decommutated data archive
ESP Experimental Solar Panel

FITS Flexible Image Transport System

FOV Field Of View FPA Focal Plane Assembly

FPGA Field Programmable Gate Arrays

GPS Global Positioning System

HK Housekeeping

IOS Instrument Operations Sheet

LED Light Emitting Diode
LYRA LYman alpha RAdiometer

LYTMR LYRA Telemetry Reformatter (software module of P2SC)
LYEDG LYRA Engineering Data Generator (software module of P2SC)

MCPM Mass Memory, Compression and Packetisation Module

MOC Mission Operation Center NDR Non Destructive Readout

OBSW On board Software
PI Principal Investigator
P2SC PROBA2 Science Center
ROB Royal Observatory of Belgium

SAA South Atlantic Anomaly
SEU Single Event Upset

SoFAST | Solar Feature Automated Search Tool

SWAP Sun Watcher using APS detector and image Processing

SWAVINT | SWAP AVerage INTensity

SWBSDG | SWAP Base Science Data Generator

SWEDG SWAP Engineering Data Generator (software module of P2SC)
SWTMR SWAP Telemetry Reformatter (software module of P2SC)

TBC To Be Confirmed

TBD To Be Defined TC Telecommand

UTC Coordinated Universal Time

UV Ultraviolet

VFC Voltage to Frequency Converter

# 7. APPENDIX Solar Activity Definitions

In the science section we use the following solar activity standards.

The standard scale for solar activity is:

- very low (almost no flares, only B)
- low (a few C flares)
- moderate (many C flares and at least an M flare)
- high (several M flares and an X flare)
- very high (continuous background of C flares, numerous M flares, more than one X flare)