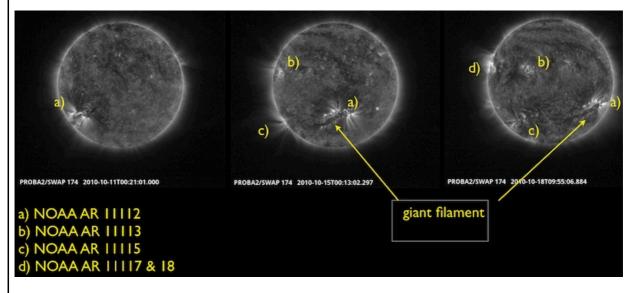
P2SC-ROB-WR-031- 20101011 Weekly report #031	P2SC Weekly report	* **** ****
Period covered: Date: Written by: Released by:	Wed Oct 27 2010 David Berghmans	Royal Observatory of Belgium PROBA2 Science Center
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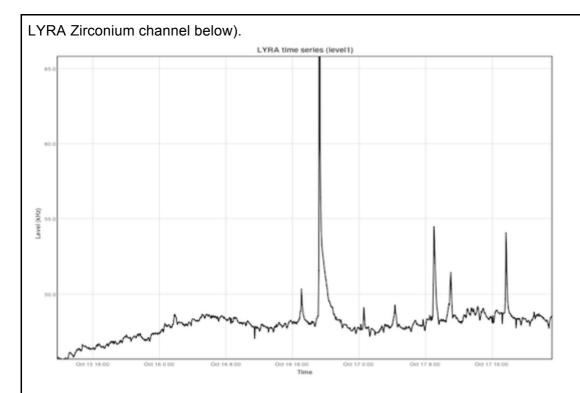
1. Science

Solar & Space weather events

Solar activity gradually increased over the period. The background GOES level increased from A6.4 in the beginning of the period to B1.0 at the end of the period. Similarly, the estimated international sunspot index (EISN) increased from 8 to 37 over the period. This increase was clearly visible in SWAP images by the popping up of new active regions.



In the second half of the week NOAA AR 11112 developed an opposite polarity island in the left handside of the active region and, as a consequence, started microflaring. This activity increase culminated in an M2.9 flare on Oct 16, 19:12UT, followed by a few more C-flares (see



The flare kernel produced an X-type diffraction pattern in the SWAP image of 19:13:04. SWAP also revealed an eruption (difference movie at 19:15.04) and a small dimming to the East of the active region. Nevertheless, the large filament in the vicinity of the active region did not erupt and SOHO/LASCO and STEREO/SECCHI didn't show any evidence of an associated, strong CME.

Scientific campaigns

None.

Outreach, papers, presentations, etc.

Julia Shugay (PROBA2 guest investigator) presented her research plan on Oct 11 11:00 in the P2SC room. It involves detecting coronal holes in SWAP images by intensity thresholding based on this, predicting the solar wind speed at L1 with the help of neural network techniques.

Also Dr. Kariyappa is still visiting the P2SC as Guest Investigators for LYRA.

To be explored

The diffraction pattern in the SWAP image at a flare kernel was a first. The M2.9 flare was the biggest flare of the present solar cycle (together with another M2.9 flare on February 6 2010).

2. LYRA instrument status

Calibration

LREP2+ LREP03 (unit 2) was scheduled (IOS00092) on Oct 14, 14:00.

Ingolf Dammasch compared the impact of the Feb 6 M2.9 flare and the Oct 16 M2.9 flare on the LYRA Al and Zr signals of unit 2. Both flares cause a similar increase in LYRA signal, for both channels.

This is quite remarkable as the Aluminium channel (and also Ly-a and Hz) has degraded considerably, due to a polymerization layer (outgassing material polymerized by UV light) on the filters. The hypothesis that this layer is blocking primarily the longer wavelenghts seems to be confirmed: the "dirt film" on the filters has very little effect on the Soft X-Ray radiation of the flare.

This is good news for the lifetime of the LYRA channels as flare detectors. On the other hand, it will make the calibration - removing the degradation effect - even more complicated.

IOS & operations

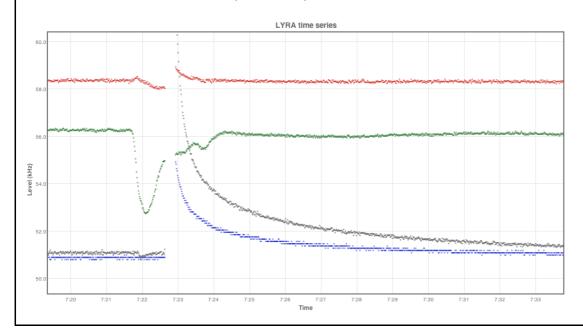
The SWAP bake-out campaign (Oct 12 10:20 - Oct 13 10:20) involved an off-pointing. LYRA was scheduled to observe 1 hour into the off-pointing period (up till Oct 12 10:20) and then go to IDLE mode. It was interesting to note that in this 1 hour off-pointing period, LYRA produced a similar pattern as during the flight mode/anti flight mode campaign.

LYRA IOS 00090 scheduled the start-up following the SWAP bake-out period from Oct 13 13:30 onwards. However, this IOS did not include a switch off the heaters and, as a consequence, the LYRA detectors became unusually hot (TEMP DD1 and TEMP DD3 \sim 47C, and TEMP DD2 \sim 50C).

To be explored

The Herzberg channel (unit 2) behaved erratically (jumps over LARs) during the second half of the week.

A pseudo-flare was observed on 2010/10/14T07:22, seen by no other solar instrument. This turned out to be an ASIC reload (see below):



3. SWAP instrument status

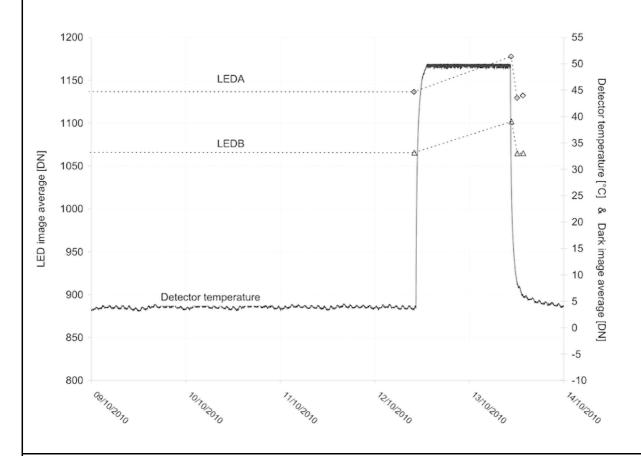
MCPM errors

The number of MCPM recoverable errors stayed fixed at 203 and the number of MCPM unrecoverable errors stayed fixed at 0.

IOS & operations

SWAP run a 24h bake-out campaign from Oct 12 10:20 onwards, preceded and followed by usual calibration sequence (IOS00186/IOS00187). During the bake-out, the SWAP detector was brought to 50C.

The comparison of the pre- and post-bake out LED sequences (by JP Halain CSL) is not easy as the temperature before and after the bake-out was different and the LEDs are known to emit less light with increasing temperature. In any case the effect of the bake-out has not been dramatic. Further LED sequences in the near future will potentially reveal more subtle trends.



SWAP detector temperature

The SWAP Cold Finger Temperature increased from 3.8C in the beginning of the period to 4.5C at the end of the period.

4. PROBA2 Science Center Status

David Berghmans was operator during this week. SWAP daily movies were created manually.

Various P2SC malfunctionings happened on October 13 late as a consequence of local network problems and a power failure. Early October 14, all pending problems were cleaned up and everything was back to nominal.

The following tools were updated on the operational server:

Software name	Update	Date	Comment
ADB	3773	Oct 11	The PHP interface of the Ancillary Data Browser was completely rewritten by C.Cabanas. Improved SQL queries give faster response.
ODP/PP_PLOT	3790, 3792	Oct 14, 15	Bug correction by E. D'Huys in the

5. Data reception & discussions with MOC

Passes

 "Packet CRC does not validate" for at least 1 BINSWAP image in pass 2630 (Oct 12), pass 2639 (Oct 13), pass 2659 (Oct 15), pass 2675 (Oct 16),

Data coverage HK

Housekeeping data is complete except for a small gap on Oct 15 around 21:34.
 Reason is unknown.

Data coverage SWAP

- The SWAP data coverage was interrupted by the SWAP bake-out campaign (see above).
- Statistics for complete week:

Total number of images between 2010 Oct 11 0UT and 2010 Oct 18 0UT: 4435

Highest cadence in this period: 30 seconds

Average cadence in this period: 136.37 seconds

Number of image gaps larger than 300 seconds: 10

Largest data gap: 1454.98 minutes

Data coverage LYRA

The LYRA STD FITS files are complete except for the known interruptions (SWAP bake-out campaign and LYRA calibration campaign).

6. APPENDIX Frequently used acronyms

ADP Ancillary Data Processor

ADPMS Advanced Data and Power Management System

AOCS Attitude and Orbit Control System

APS Active Pixel image Sensor

ASIC Application Specific Integrated Circuit

BBE Base Band Equipment CME Coronal Mass Ejection

COGEX Cool Gas Generator Experiment
CRC Cyclic Redundancy Check
DR Destructive Readout

DSLP Dual Segmented Langmuir Probe
EIT Extreme ultraviolet Imaging Telescope
FITS Flexible Image Transport System

FOV Field Of View FPA Focal Plane Assembly

FPGA Field Programmable Gate Arrays

GPS Global Positioning System
HAS High Accuracy Star tracker

HK Housekeeping

ICD Interface Control Document
IIU Instrument Interface Unit
IOS Instrument Operations Sheet

LED Light Emitting Diode LEO Low Earth Orbit

LYRA LYman alpha RAdiometer

LYTMR LYRA Telemetry Reformatter (software module of P2SC)
LYEDG LYRA Engineering Data Generator (software module of P2SC)

MCPM Mass Memory, Compression and Packetisation Module

MOC
NDR
OBET
OBSW
PE
Mission Operation Center
Non Destructive Readout
On board Elapsed Time
On board Software
Proximity Electronics

PGA Programmable Gain Amplifier

PI Principal Investigator P2SC PROBA2 Science Center

PPT Pointing, Positioning and Time (software module of P2SC)

ROB Royal Observatory of Belgium SAA South Atlantic Anomaly Spacecraft Operation System

SEU | Single Event Upset

SOHO Solar and Heliospheric Observatory

SWAP Sun Watcher using APS detector and image Processing

SWBSDG SWAP Base Science Data Generator

SWEDG SWAP Engineering Data Generator (software module of P2SC) SWTMR SWAP Telemetry Reformatter (software module of P2SC)

TBC
TBD
To Be Confirmed
To Be Defined
To Be Written
TC
Telecommand

TPMU Thermal Plasma Measurement Unit

UTC Coordinated Universal Time

UV Ultraviolet